

Search for pulsed γ -ray emission from globular cluster M28

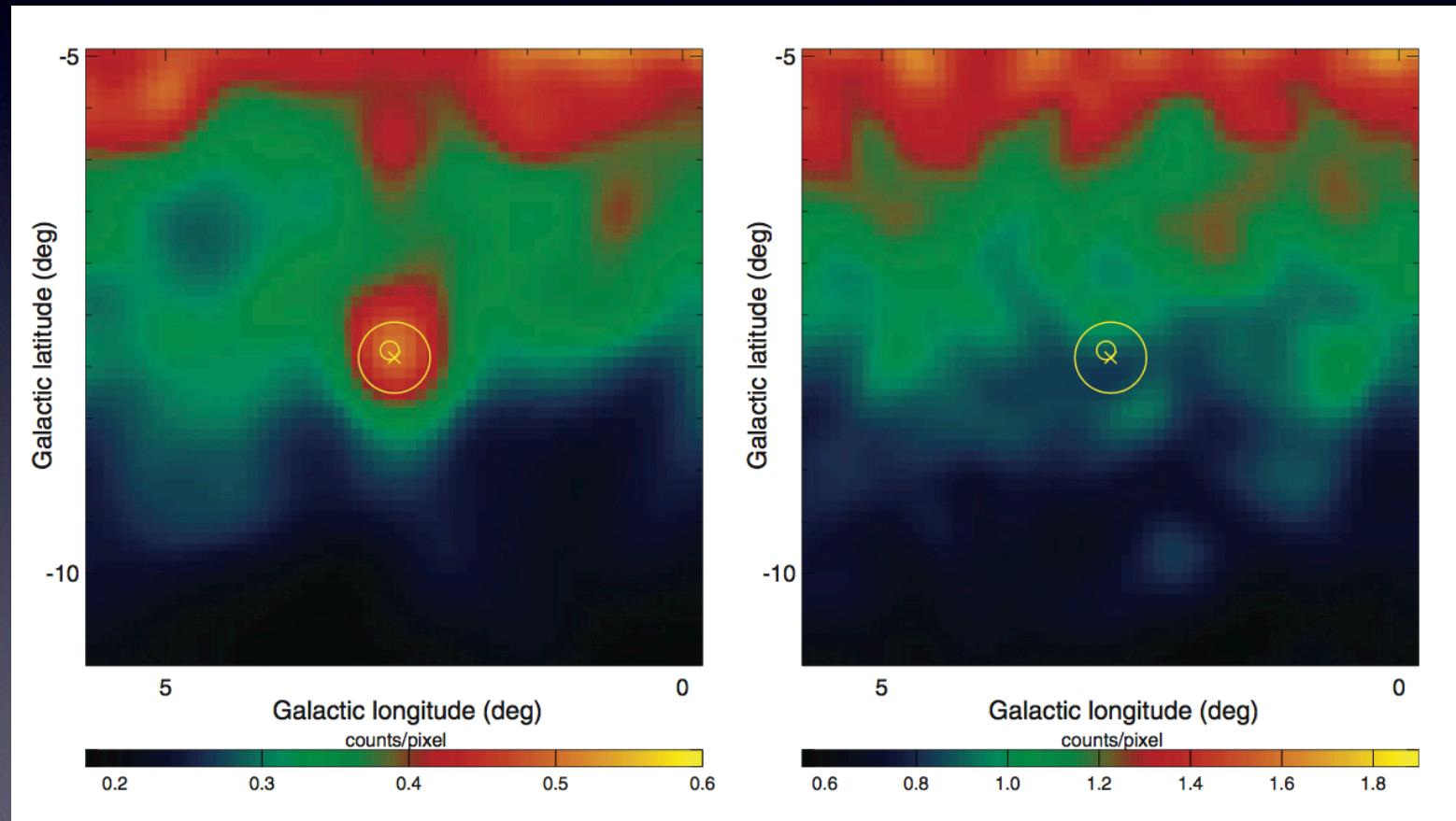
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K. S. Cheng (HKU)

submitted to ApJL

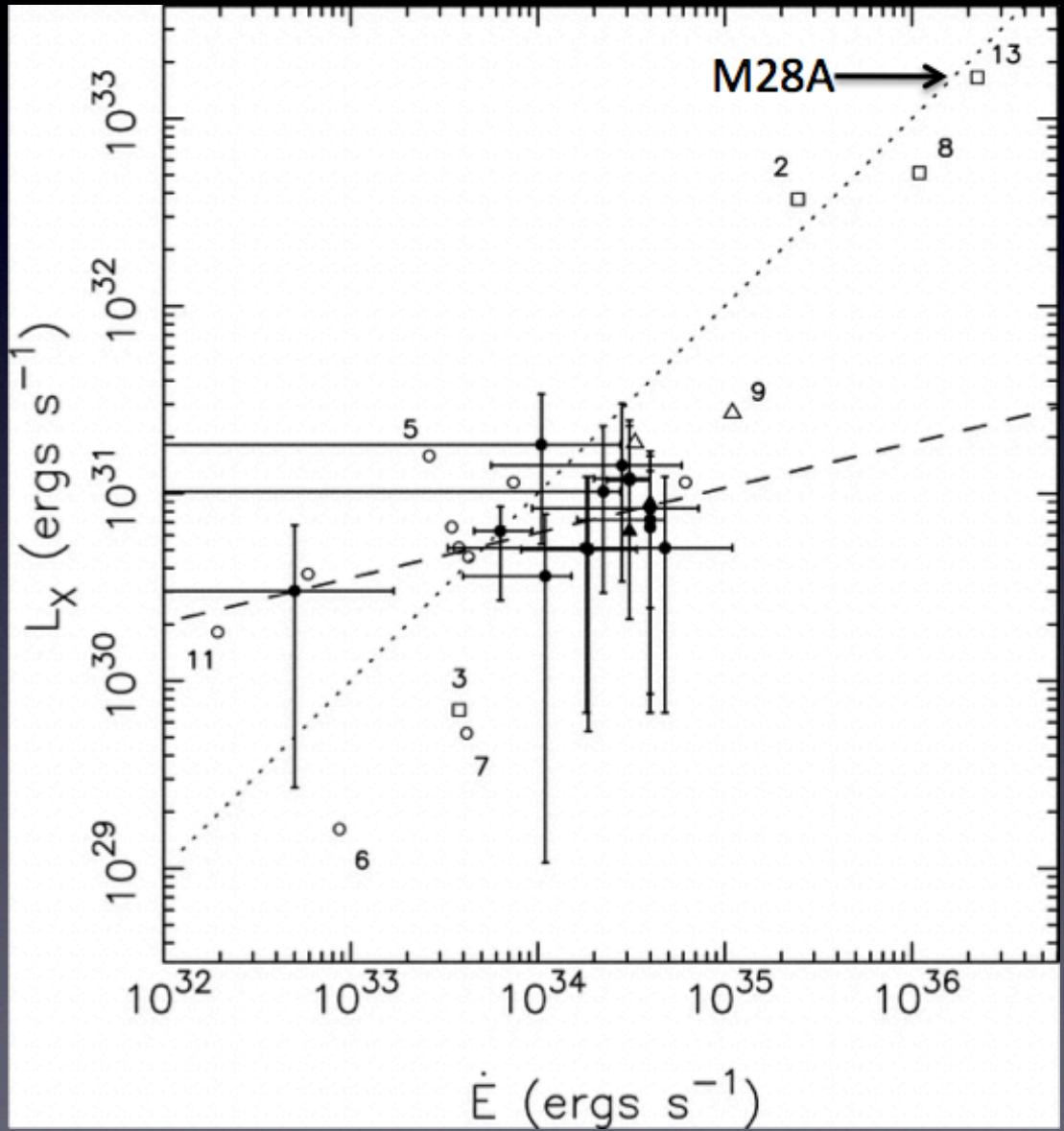
Pulsed γ -ray from Globular Clusters

PSR J1823-3021A @ NGC6624



Freire et al. (2011)

M28A is so energetic



$dE/dt = 2.2 \times 10^{36} \text{ erg/s}$

Strongest MSP in GC

Bogdanov et al. (2011)

Starting point

PSR J1824-2452

RAJ 18:24:32.00790550 0 0.00001054

DECJ -24:52:10.8076448 0 0.0020416

F0 327.4056060517495439 0 0.0000000000233327

F1 -1.735361869603E-13 0 1.123466933817E-18

PEPOCH 53800.000000

DM 0.000356

EPHEM DE405

TZRMJD 53859.26176012789721

TZRSITE 1

JUMP_1 -0.00015523 1 0.00000010

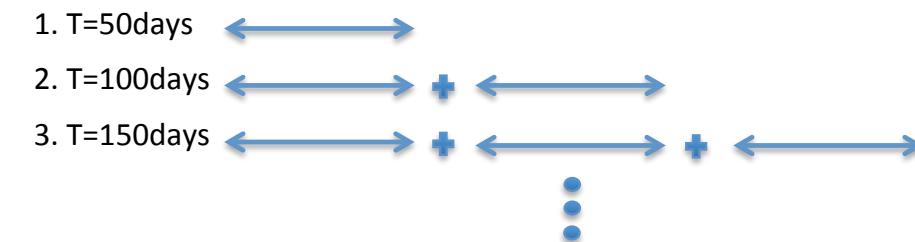
JUMP_2 0.00136886 1 0.00000012

JUMP_3 0.00129312 1 0.00000061

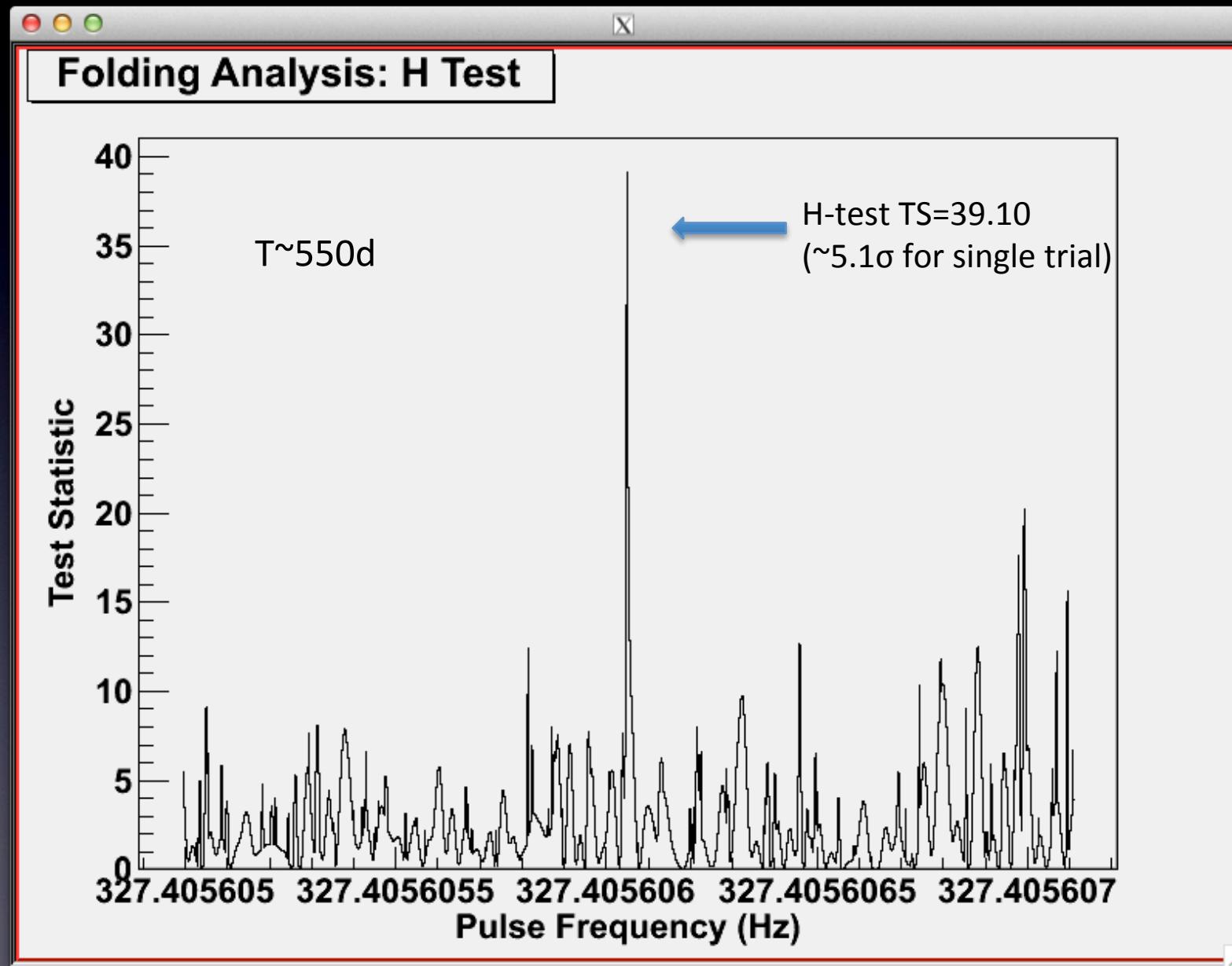
Ray et al. 2008

Pulsation search

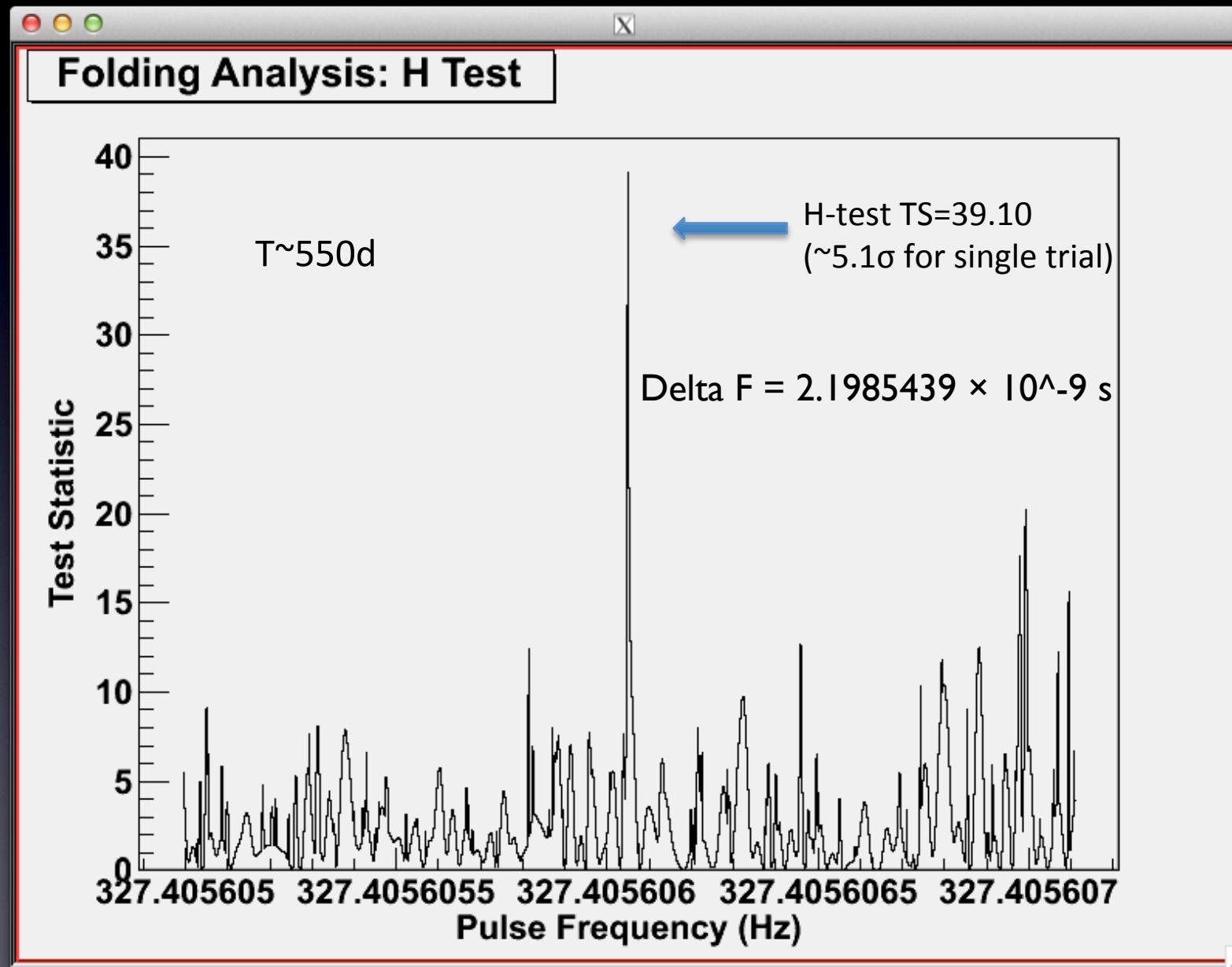
- By fixing the first period derivative (f_1) at it's best-fit value, we search for coherent signal by stacking up photons with 50 days interval



Pulsation search



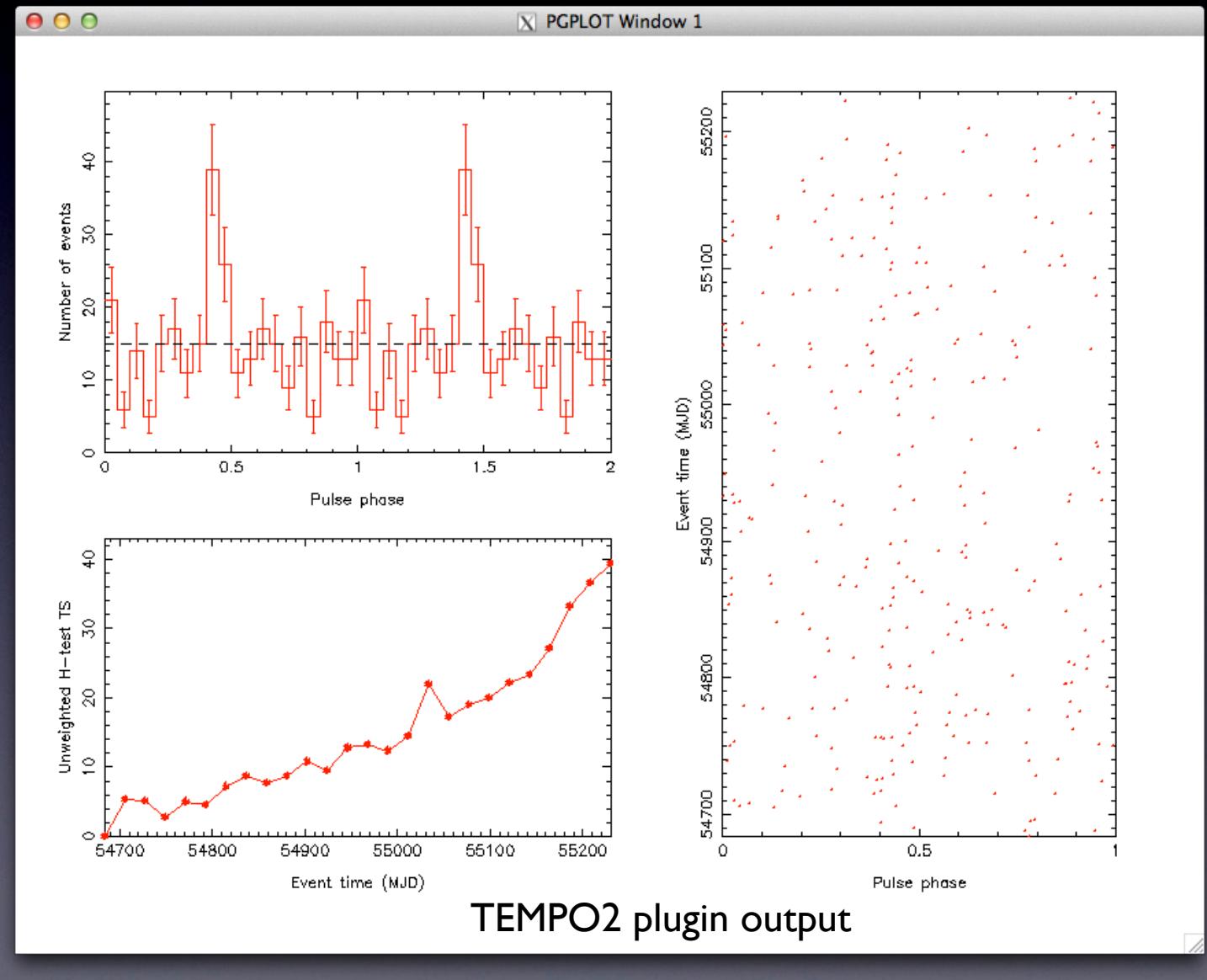
Pulsation search



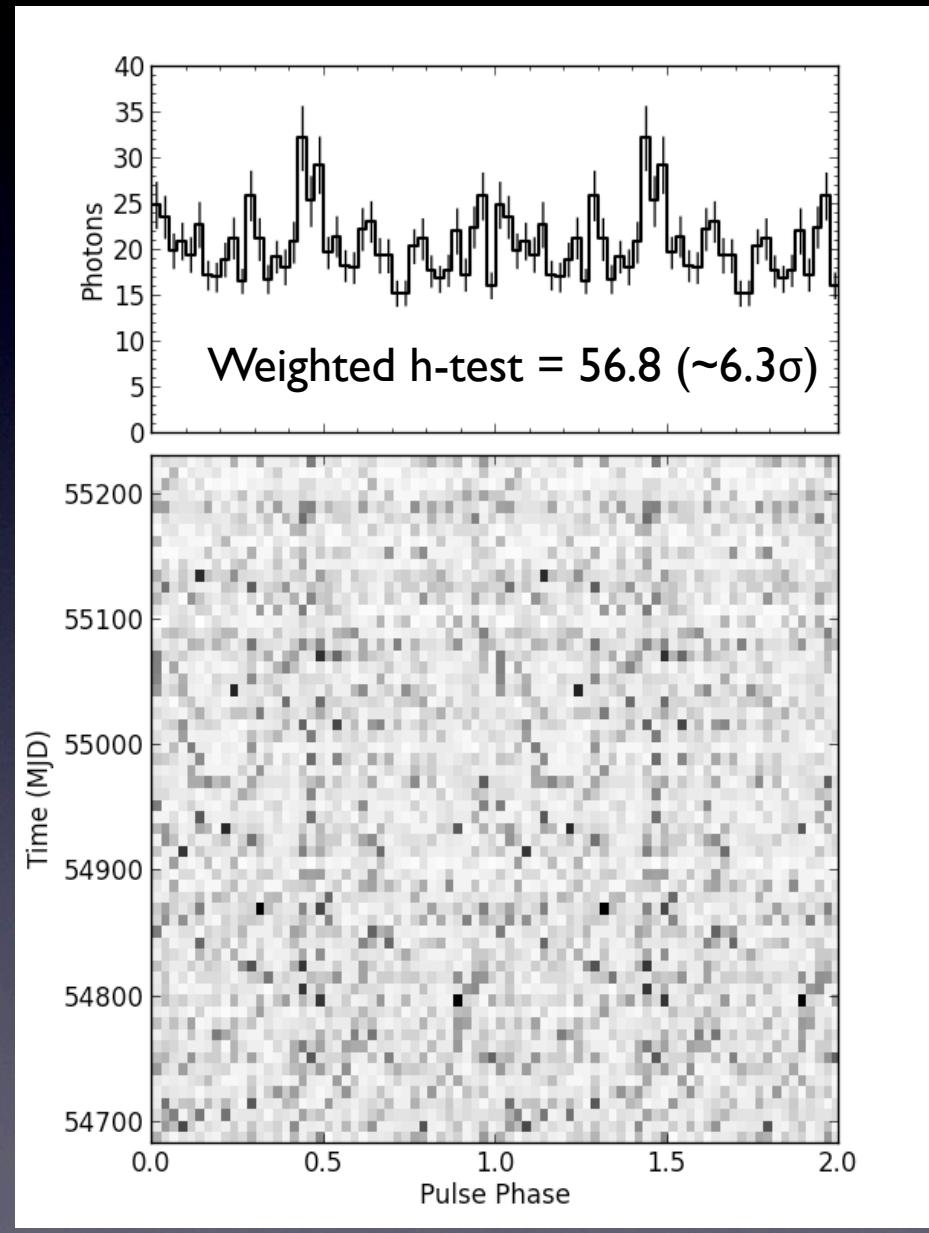
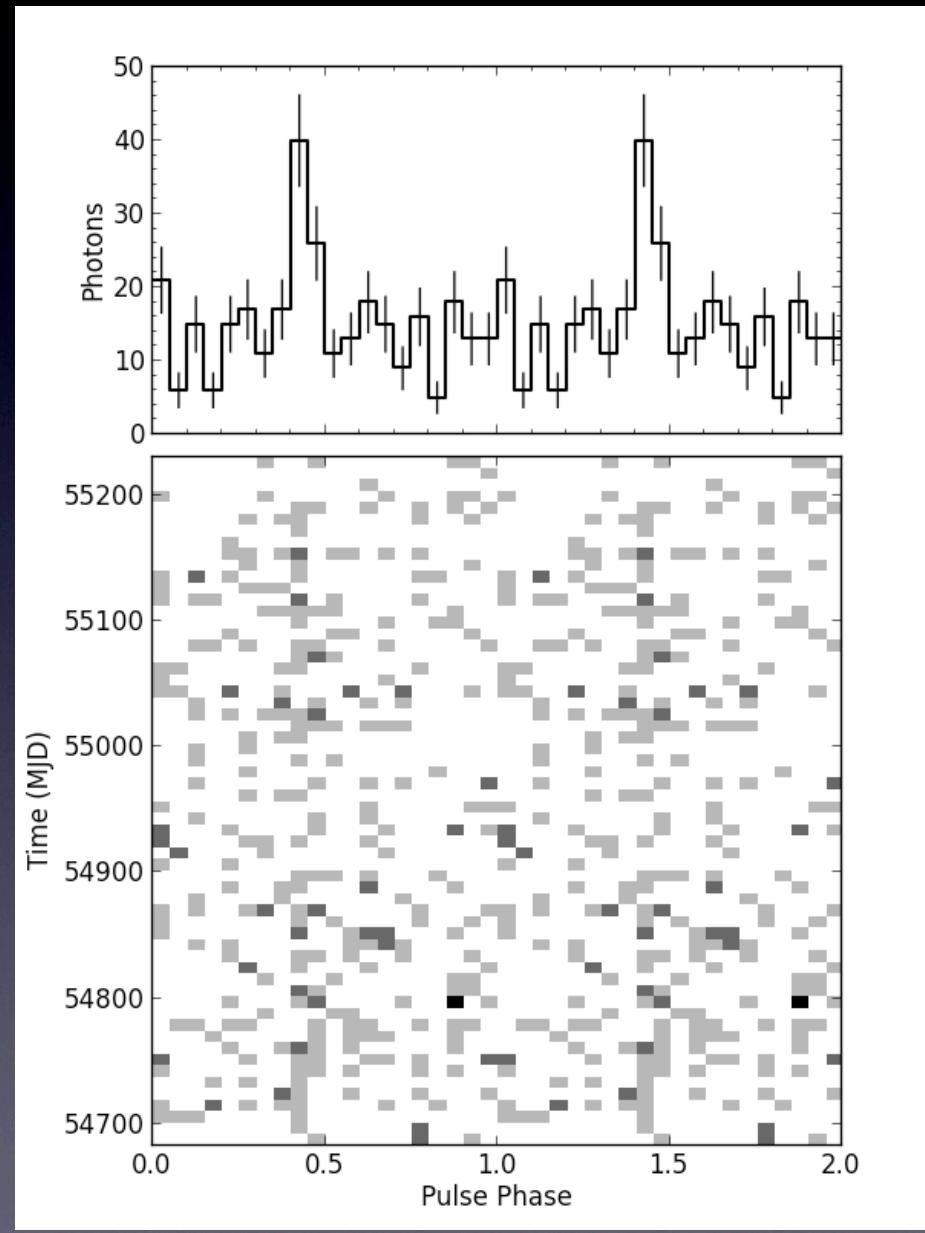
Folded lightcurve

ROI = 0.5deg

E_min = 720 MeV

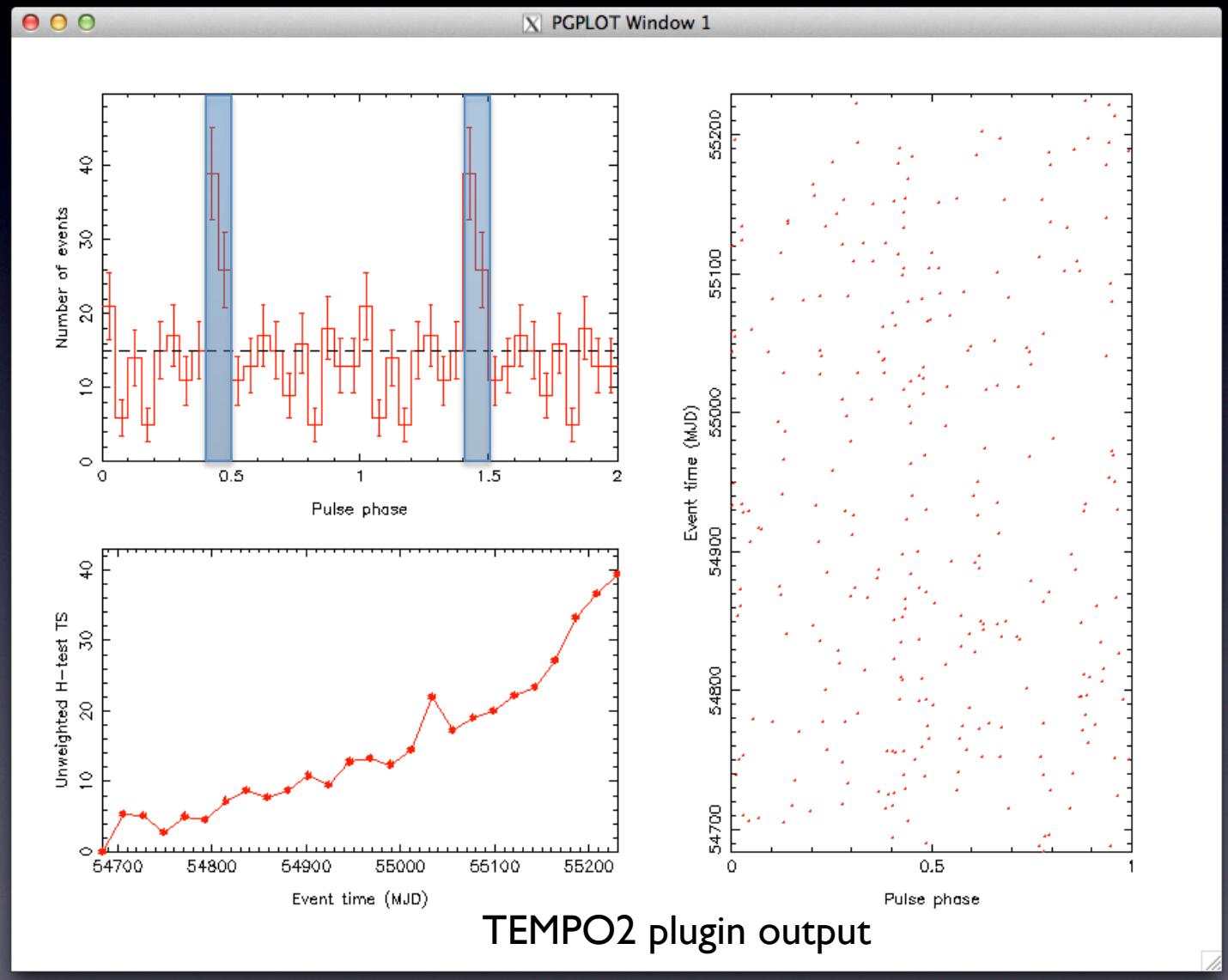


Weighted photon

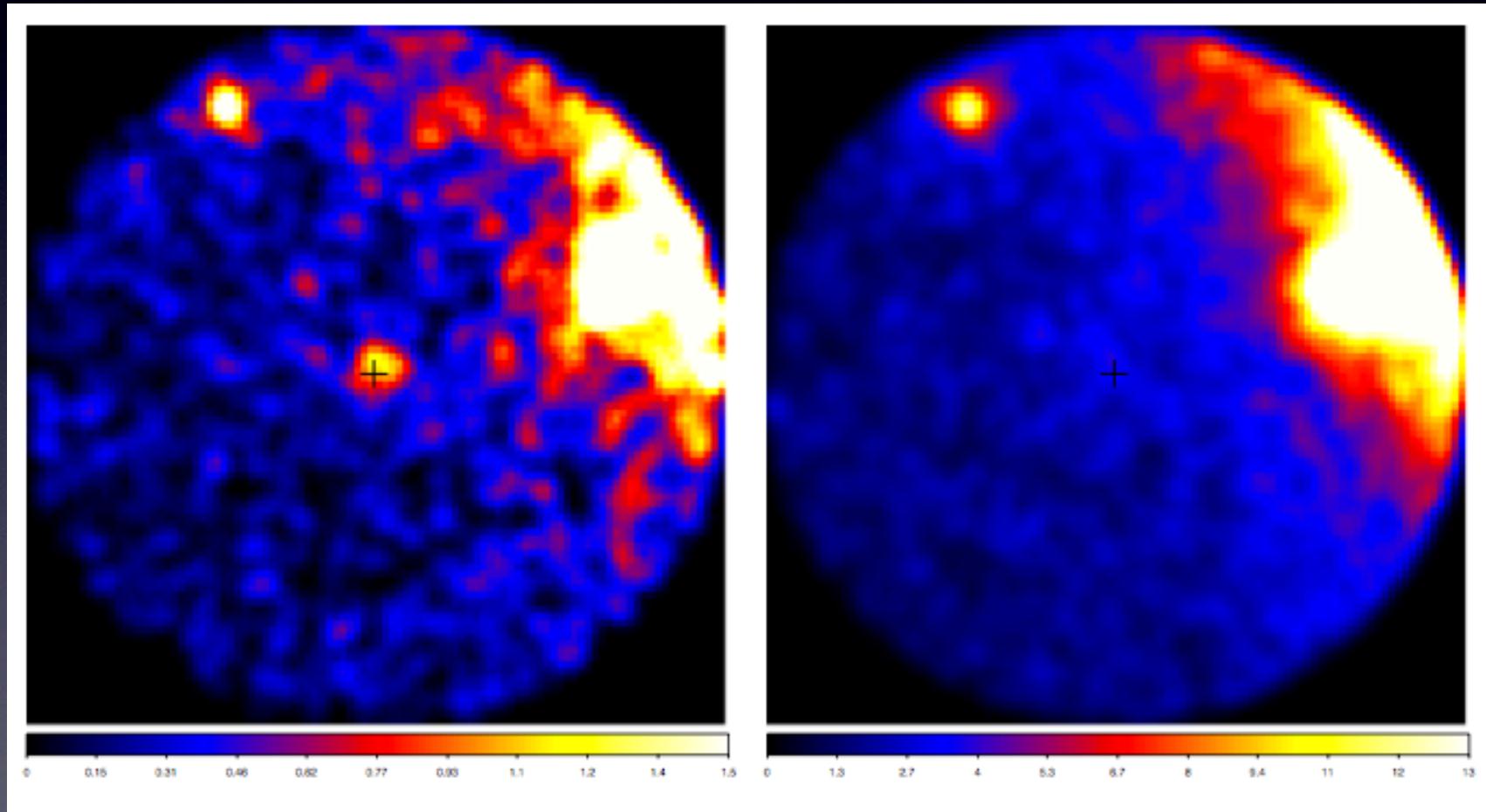


Phase resolved analysis

We defined phases between 0.4-0.5 as onpulse interval



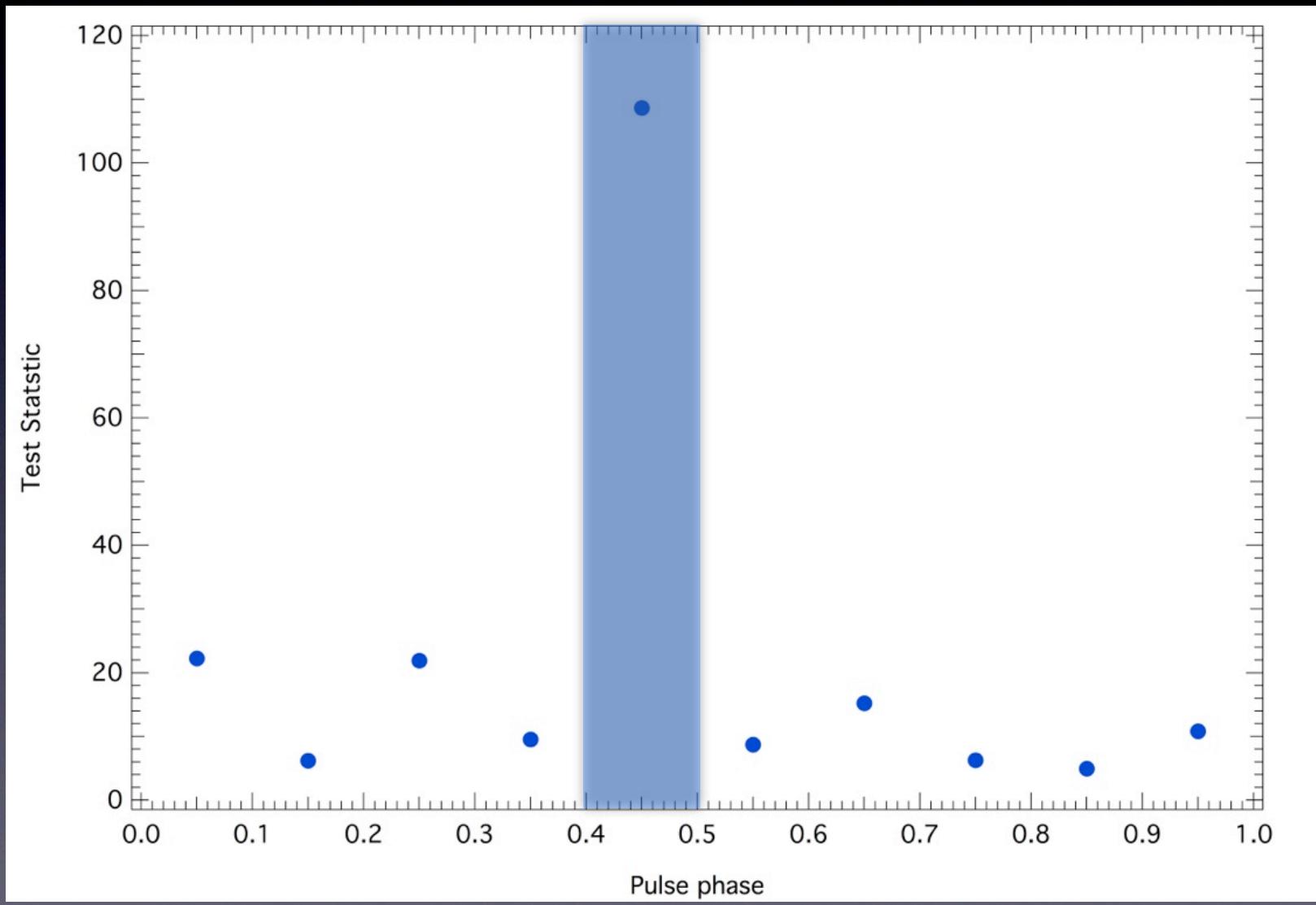
Phase resolved analysis



Pulse on ($\Phi=0.4-0.5$)

Pulse off ($\Phi=0.5-1.4$)

Phase resolved analysis



END